

Mueller MT2 Problem 1

a. In one hour, a mass of

$$(1.1 \times 10^6) (1.1 \times 10^3 \text{ kg/m}^3) = 1.1 \times 10^9 \text{ kg}$$

of water falls. This generates

$$.85, ((1.1 \times 10^9 \text{ kg}) (9.8 \text{ m/s}^2) (350 \text{ m})) = 3.20 \times 10^{12} \text{ Joules}$$

of energy (taking efficiency into account).

Finally, the power is

$$\frac{\Delta E}{\Delta t} = \frac{3.20 \times 10^{12} \text{ J}}{3600 \text{ s}} = 8.9 \times 10^8 \text{ W.} \quad 2 \text{ points}$$

b. It takes

$$(1.1 \times 10^9 \text{ kg}) (9.8 \text{ m/s}^2) (350 \text{ m}) = 3.77 \times 10^{12} \text{ J}$$

of energy to pump the water up. the

plant consumes $150 \times 10^6 \text{ W}$ of power, but

only uses

$$.85, (150 \times 10^6 \text{ W}) = 130 \text{ W}$$

so it takes

$$\Delta t = \frac{\Delta E}{P} = \frac{3.77 \times 10^{12} \text{ J}}{130 \text{ W}} = 2.9 \times 10^4 \text{ s} = 8.24 \text{ hours}$$

c. Efficiency is given as

$$\eta = \frac{\text{Energy out}}{\text{Energy in}} = (.85) (.85) \frac{\text{Energy in}}{\text{Energy in}} \approx .72$$

Alternatively, the energy put in is

$$(150 \times 10^6 \text{ W}) (2.9 \times 10^4 \text{ s}) = 4.4 \times 10^{12} \text{ J}$$

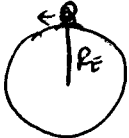
The energy out, from part (a) is $3.20 \times 10^{12} \text{ J}$,

so

$$\eta = \frac{3.2 \times 10^{12} \text{ J}}{4.4 \times 10^{12} \text{ J}} \approx .72$$

Problem 2 Mueller Midterm 2 Spring 2010

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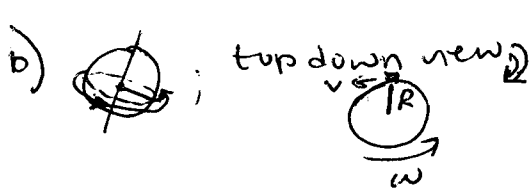
a)  $R_E = 7 \times 10^3 \text{ km}, g = 9.8 \text{ m/s}^2$

For a satellite orbiting just above ground,

$\Sigma F_c = ma_c = \frac{mv^2}{R} = mg$. (Gravitational acceleration is the centripetal acceleration, and near earth's surface $a_{grav} = g$.)

$$\Rightarrow v = \sqrt{Rg} = \sqrt{(7 \times 10^6 \text{ m})(9.8 \text{ m/s}^2)} = \underline{8282.5 \text{ m/s}}$$

The kinetic energy of the satellite is $\frac{1}{2}mv^2 = \frac{1}{2}mgR = \underline{m \cdot 3.43 \times 10^7 \text{ J}}$



The velocity of the rocket relative to the center of the earth when it is just sitting on the surface is found with $v' = R\omega$, where

ω is the angular velocity of earth. We can obtain ω by using the period of earth's rotation, which is one day:

$$\omega = 2\pi/T = 2\pi/24 \cdot 60 \cdot 60 = 7.27 \times 10^{-5} \text{ s}^{-1}$$

Then the velocity (magnitude) of the rocket relative to the center of the earth, when the rocket is sitting at the equator, is $v' = R\omega$

$$= (7 \times 10^6 \text{ m})(7.27 \times 10^{-5} \text{ s}^{-1}) = \underline{509 \text{ m/s}}$$
, which is the ~~amount~~

the necessary velocity is reduced (for going into orbit).

Relative to the launch pad, the remaining velocity needed to enter an orbit close to earth's surface is $\Delta v = |v' - v| = |R\omega - \sqrt{Rg}| = \underline{7774 \text{ m/s}}$

The kinetic energy of the satellite relative to the launch pad is

$$\frac{1}{2}m(R\omega - \sqrt{Rg})^2 = \frac{1}{2}m(R^2\omega^2 - 2R\omega\sqrt{Rg} + Rg) = \underline{3.022 \times 10^7 \text{ m} \cdot \text{J}}$$

Then the necessary kinetic energy for orbit is reduced by

$$(\text{original}) - (\text{new}) = \cancel{\frac{1}{2}mRg} - \frac{1}{2}mR^2\omega^2 + mR\omega\sqrt{Rg} - \cancel{\frac{1}{2}mRg} = \underline{m \cdot 4.09 \times 10^6 \text{ J}}$$

Mueller Problem 2

(2)

As a percentage of the energy required in part a, the kinetic energy is reduced by $\frac{\Delta KE}{KE_{\text{orig}}} \times 100 = 100 \times \frac{4.09 \times 10^6}{3.43 \times 10^7} = 12\%$.

c) $g_m = \frac{1}{6} g_E$; $\rho_m = \rho_E$. We want r_m .

$$\rho_m = \rho_E \Rightarrow \frac{M_m}{R_m^3} = \frac{M_E}{R_E^3} \Rightarrow M_E = M_m \frac{R_E^3}{R_m^3}$$

The force acting on some mass m at the surface of each is

random little mass

random little mass

moon (subscript)

$$\textcircled{1} \frac{G m M_E}{R_E^2} = m g_E, \quad \textcircled{2} \frac{G m M_m}{R_m^2} = m g_m$$

Divide

$$\frac{\textcircled{1}}{\textcircled{2}} = \frac{\frac{G M_E}{R_E^2}}{\frac{G M_m}{R_m^2}} = \frac{M_E g_E}{M_m g_m} = 6$$

$$\Rightarrow \frac{M_m R_E^3 / R_m^3}{R_E^2 M_m / R_m^3} \Rightarrow R_E = 6 R_m \Rightarrow R_m = \underline{1.167 \times 10^6 \text{ m}}$$

Lecture 3

3.) a.)



$$L = 1.100 \text{ m}$$

$$d = 1.021 \text{ m}$$

$$M = 5.000 \text{ kg}$$

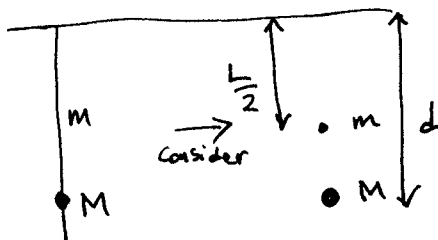
$$m = 1.000 \text{ kg}$$

$$I_P = \frac{1}{12} ML^2 + M\left(\frac{L}{2}\right)^2 + \frac{Md^2}{2}$$

$$I_P = \frac{1}{3} mL^2 + Md^2$$

$$I_P = \frac{5.616 \text{ kgm}^2}{+1}$$

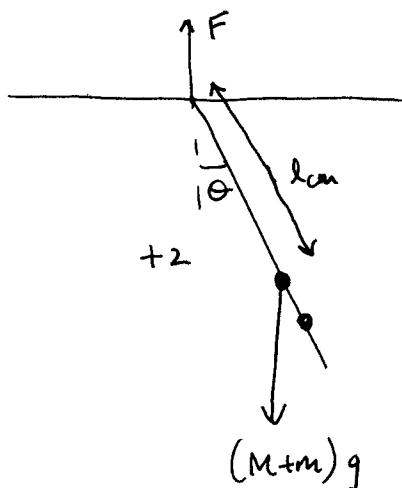
b.)



$$l_{cm} = \frac{m \frac{L}{2} + Md}{m + M}$$

$$l_{cm} = \frac{0.9425 \text{ m}}{+1}$$

c.)



$$\tau = -(M+m)g l_{cm} \sin \theta$$

$$\tau \approx -\frac{(m \frac{L}{2} + Md)g \theta}{+3} \approx 5.655 g \theta \approx 55.4 \theta$$

$$\tau = I \alpha = I \frac{d^2 \theta}{dt^2}$$

$$d.) \quad \frac{d^2 \theta}{dt^2} = \frac{\tau}{I} = -\frac{(m \frac{L}{2} + Md)g}{\frac{1}{3} mL^2 + Md^2} \theta$$

$$\text{Take } \theta(t) = A \sin\left(\frac{2\pi t}{T}\right)$$

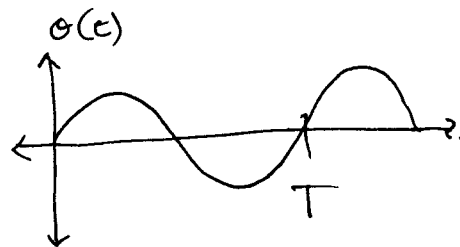
$$\frac{d\theta}{dt} = \frac{2\pi}{T} A \cos\left(\frac{2\pi t}{T}\right)$$

$$\frac{d^2 \theta}{dt^2} = -\left(\frac{2\pi}{T}\right)^2 A \sin\left(\frac{2\pi t}{T}\right)$$

$$\frac{d^2 \theta}{dt^2} = -\left(\frac{2\pi}{T}\right)^2 \theta$$

$$\rightarrow \left(\frac{2\pi}{T}\right)^2 =$$

$$\frac{m \frac{L}{2} + Md}{\frac{1}{3} mL^2 + Md^2} g \rightarrow T = 2\pi \sqrt{\frac{\frac{1}{3} mL^2 + Md^2}{(m \frac{L}{2} + Md)g}}$$



T is the period

$$T = \frac{2.000 \text{ sec}}{2.000}$$

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Mueller

a) Since power is constant

$$\frac{1}{2}mv^2 = KE = P \cdot t$$

$$t = \frac{\frac{1}{2}mv^2}{P} = 3.7s$$

b) Looking at the front of the SUV:



Both the centrifugal force and gravity can be taken at the C.o.M.

$$F_c = \frac{mv^2}{r}$$

$$|\tau_c| = |\vec{R} \times \vec{F}_c| = \frac{mv^2}{r} h_{cm} = \frac{2.0 \times 10^6 \text{ kg m}^3/\text{s}^2}{r}$$

c) It begins to tip over when

$$mg \cdot \frac{w}{2} = |\tau_g| = |\tau_c| = \frac{mv^2}{r} h_{cm}$$

$$r = \frac{2v^2 h_{cm}}{gw} = 68m$$

a)

$$V_{CM,x} = \frac{M V_M}{M+m} = \frac{40}{3} \text{ mph}$$

$$V_{CM,y} = \frac{m V_m}{M+m} = \frac{20}{3} \text{ mph}$$

No external forces

 $\Rightarrow \vec{V}_{CM}$ same before and after collision

b) Before collision:

- Rest frame: $KE_{tot} = \frac{1}{2} M V_M^2 + \frac{1}{2} m V_m^2$

$$= \frac{1}{2} (2400 \text{ kg}) \left(\frac{80}{9} \frac{\text{m}}{\text{s}} \right)^2 + \frac{1}{2} (1200 \text{ kg}) \left(\frac{80}{9} \frac{\text{m}}{\text{s}} \right)^2$$

$$= \boxed{142222 \text{ J}}$$

- CM frame: $KE_{tot} = \frac{1}{2} M \left[\left(20 - \frac{40}{3} \text{ mph} \right)^2 + \left(\frac{20}{3} \text{ mph} \right)^2 \right]$

$$+ \frac{1}{2} m \left[\left(\frac{40}{3} \text{ mph} \right)^2 + \left(20 - \frac{20}{3} \right)^2 \right]$$

$$= \boxed{63210 \text{ J}}$$

After collision:

- Rest frame: $KE_{tot} = \frac{1}{2} M V_{Mf}^2 + \frac{1}{2} m V_{mf}^2$

$$= \frac{1}{2} M \left[\left(\frac{40}{3} \text{ mph} \right)^2 + \left(\frac{20}{3} \text{ mph} \right)^2 \right]$$

$$+ \frac{1}{2} m \left[\left(\frac{40}{3} \text{ mph} \right)^2 + \left(\frac{20}{3} \text{ mph} \right)^2 \right]$$

$$= \boxed{79012 \text{ J}}$$

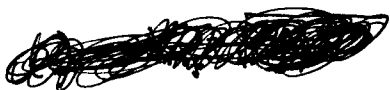
- CM frame:

Both cars move with CM frame after collision ≈ 0

$$KE_{tot} = 0$$

c)

$$\vec{V}_M = \vec{V}_m = \left(\frac{40}{3} \text{ mph}, \frac{20}{3} \text{ mph} \right)$$



5] cont.

Mueller
MT2

$$d) a_{M,x} = \frac{\Delta V_{Mx}}{50 \text{ ms}} = \frac{\left(\frac{40}{3} - 20\right) \text{ mph}}{50 \text{ ms}} = \frac{\left(-\frac{20}{3} \text{ mph}\right)}{50 \text{ ms}} = -59.26 \frac{\text{m}}{\text{s}^2}$$

$$a_{M,y} = \frac{\Delta V_{My}}{50 \text{ ms}} = \frac{\frac{20}{3} \text{ mph}}{50 \text{ ms}} = 59.26 \frac{\text{m}}{\text{s}^2}$$

$$\Rightarrow a_M = \sqrt{(-59.26)^2 + (59.26)^2} = \boxed{83.8 \frac{\text{m}}{\text{s}^2}}$$

$$a_{m,x} = \frac{\Delta V_{m,x}}{50 \text{ ms}} = \frac{\frac{40}{3} \text{ mph}}{50 \text{ ms}} = 118.52 \frac{\text{m}}{\text{s}^2}$$

$$a_{m,y} = \frac{\Delta V_{m,y}}{50 \text{ ms}} = \frac{\left(\frac{20}{3} - 20\right) \text{ mph}}{50 \text{ ms}} = \frac{-\frac{40}{3} \text{ mph}}{50 \text{ ms}} = -118.52 \frac{\text{m}}{\text{s}^2}$$

$$a_m = \sqrt{(118.52)^2 + (-118.52)^2} = \boxed{167.6 \frac{\text{m}}{\text{s}^2}}$$

Note: $F_{M \rightarrow m} = F_{m \rightarrow M}$ so $M a_M = m a_m$

$$\Rightarrow \frac{a_m}{a_M} = \frac{M}{m} = 2$$

6. a. The power P is converted into kinetic energy $K(t)$, so $P = \frac{dK}{dt}$

then $K = \int dK = \int P dt = P \cdot t$ (since P is const.)

$$K(t) = \frac{1}{2} m (v(t))^2 = P \cdot t$$

$$\Rightarrow v(t) = \sqrt{\frac{2Pt}{m}}$$

Note: $a = \frac{dv}{dt}$ is not constant!
This means you cannot use $x = \frac{1}{2} at^2$!

Displacement $x(t) = \int_0^t v(t') dt' = \sqrt{\frac{2P}{m}} \int_0^t t'^{1/2} dt' = \frac{2}{3} \sqrt{\frac{2P}{m}} t^{3/2}$

Let D be distance between stop signs and T be the time to travel between them. Then:

$$x(T) = D = \sqrt{\frac{8P}{9m}} T^{3/2} \rightarrow \text{solve for } T$$

$$T = D^{2/3} \left(\frac{9m}{8P}\right)^{1/3} \text{ for each stop sign}$$

$$1 \text{ mile} = \underbrace{(6 \times 250 \text{ meters})}_D + \underbrace{109.3 \text{ meters}}_{D_{\text{end}}}$$

$$T_{\text{tot}} = 6 \times T + T_{\text{end}} = \left(\frac{9m}{8P}\right)^{1/3} (6D^{2/3} + D_{\text{end}}^{2/3}) = 193 \text{ sec.} \\ = 3.2 \text{ min.}$$

6. b. Constant v means that your power output is just enough to cancel frictional losses:

$$dW = \vec{f} \cdot d\vec{x} = -f dx$$

$$P = -\frac{dW}{dt} = +f \frac{dx}{dt} = f v$$

velocity $v = \frac{1 \text{ mi}}{T_{\text{tot}}} = \frac{1609.3 \text{ m}}{193 \text{ s}} = 8.34 \text{ m/s}$
 $= 18.7 \text{ mi/hr}$

It is not strictly correct to describe the friction by a single coefficient. There is energy loss due to the kinetic friction at each moving part of the bicycle, as well as deformation of the wheels and air resistance. However, we will approximate the friction, using the effective coefficient μ , as $f = mg\mu$.

Then: $P = mg\mu v$

$$= 90 \text{ kg} * 9.8 \frac{\text{m}}{\text{s}^2} * 0.05 * 8.34 \text{ m/s}$$

$$= 368 \text{ W}$$